

265(5): Further Details of the Derivation of the Landau Precession from the Thomas Precession.

Consider the infinitesimal line element:

$$ds^2 = c^2 d\tau^2 = c^2 dt^2 - dr^2 - r^2 d\theta^2 \quad (1)$$

in plane polar coordinates, where τ is the proper time. The Thomas precession is defined as:

$$\theta' = \theta + \omega t, \quad (2)$$

$$d\theta' = d\theta + \omega dt. \quad (3)$$

where ω is an angular velocity. Therefore:

$$\begin{aligned} (d\theta')^2 &= (d\theta + \omega dt)^2 \\ &= d\theta^2 + 2\omega d\theta dt + \omega^2 dt^2 \quad (4) \end{aligned}$$

It follows that under the Thomas precession:

$$\begin{aligned} ds^2 &= c^2 d\tau^2 = c^2 dt^2 - dr^2 - r^2 (d\theta^2 + 2\omega d\theta dt + \omega^2 dt^2) \\ &= (c^2 - r^2 \omega^2) dt^2 - dr^2 - r^2 d\theta^2 - 2\omega r^2 d\theta dt \quad (5) \end{aligned}$$

The Thomas precession is defined by the velocity:

$$V_\theta = r\omega \quad (6)$$

$$\text{so } ds^2 = \left(1 - \frac{V_\theta^2}{c^2}\right) c^2 dt^2 - dr^2 - r^2 d\theta^2 - 2\omega r^2 d\theta dt \quad (7)$$

Now use: $\omega = \frac{d\theta}{dt}, \quad V_\theta = \omega r \quad (8)$

$$so \quad d\theta = \omega dt = \frac{V_\theta}{r} dt. \quad - (9)$$

It follows that:

$$ds^2 = \left(1 - \frac{3V_\theta^2}{c^2}\right) c^2 dt^2 - dr^2 - r^2 d\theta^2 \quad - (10)$$

The orbital velocity is defined by:

$$v^2 = \left(\frac{dr}{dt}\right)^2 + r^2 \left(\frac{d\theta}{dt}\right)^2, \quad - (11)$$

$$so \quad c^2 d\tau^2 = \left(1 - \frac{3V_\theta^2}{c^2}\right) c^2 dt^2 - v^2 dt^2 \quad - (12)$$

$$i.e. \quad \left(\frac{d\tau}{dt}\right)^2 = 1 - \frac{3V_\theta^2}{c^2} - \frac{v^2}{c^2} \quad - (13)$$

The Lorentz factor is modified to:

$$\gamma = \frac{dt}{d\tau} = \left(1 - \frac{3V_\theta^2}{c^2} - \frac{v^2}{c^2}\right)^{-1/2} \quad - (14)$$

$$\sim \frac{v^2}{c^2} + \frac{3V_\theta^2}{c^2}$$

$$if \quad v \ll c \text{ and } V_\theta \ll c \quad - (15)$$

3) The total linear velocity is:

$$\underline{v} = \underline{v}_r + \underline{\omega} \times \underline{r} \quad - (16)$$

where

$$\underline{v}_r = \frac{dr}{dt} \underline{e}_r \quad - (17)$$

so the total kinetic energy is:

$$\begin{aligned} T &= \frac{1}{2} m (\underline{v}_r^2 + \underline{\omega} \times \underline{r} \cdot \underline{\omega} \times \underline{r}) \\ &= \frac{1}{2} m (\underline{v}_r^2 + \omega^2 r^2) \quad - (18) \end{aligned}$$

if $\underline{\omega} \perp \underline{r} \quad - (19)$

the radial kinetic energy is:

$$T_r = \frac{1}{2} m v^2 \quad - (20)$$

and the angular kinetic energy is:

$$T_\theta = \frac{1}{2} m \omega^2 r^2 = \frac{1}{2} I \omega^2 \quad - (21)$$

where the moment of inertia is:

$$I = m r^2 \quad - (22)$$

For a constant ω :

$$T = \frac{1}{2} m \omega^2 r^2 = m \int \omega^2 r dr \quad - (23)$$

which is the integral of the centrifugal force.
In generating the Thomas precession, the only

4) velocity considered is:

$$\underline{V}_\theta = \underline{\omega} \times \underline{r} \quad - (24)$$

i.e

$$V_\theta = \omega r \underline{e}_\theta \quad - (25)$$

So

$$\underline{F} = m\omega^2 r \underline{e}_r = - \frac{\partial \mathcal{U}}{\partial r} = - \frac{mMG}{r^2} \quad - (26)$$

and

$$\int F dr = \frac{1}{2} m V_\theta^2 = \frac{mMG}{r} \quad - (27)$$

So

$$\boxed{V_\theta^2 = \frac{2mG}{r}} \quad - (28)$$

The infinitesimal line element (12) is therefore:

$$ds^2 = c^2 d\tau^2 = \left(1 - \frac{6mG}{c^2 r}\right) c^2 dt^2 - v^2 dt^2 \quad - (29)$$

The additional time dilation due to Thomas precession is:

$$\begin{aligned} \frac{dt}{d\tau} &= \left(1 - \frac{6mG}{c^2 r}\right)^{-1/2} \\ &\sim \frac{3mG}{c^2 r} \quad - (30) \end{aligned}$$

if $\frac{MG}{c^2} \ll r \quad - (31)$

The definition (24) of \dot{V}_θ used is the Thomas precession corresponds to a turning point of any orbit:

$$\underline{V_r} = \underline{0} \quad - (32)$$

Consider now the effect of the Thomas precession on an elliptical orbit:

$$r = \frac{d}{1 + \epsilon \cos \theta} \quad - (33)$$

The turning point occurs at:

$$r = d = a(1 - \epsilon^2) \quad - (34)$$

where d is the half right latitude, ϵ is the eccentricity, and a the semi major axis. The minimum value of r is eq. (33) is the perihelion, when r is closest to M . The perihelion is a turning point.

So at the perihelion:

$$\boxed{\frac{dt}{d\tau} = \frac{3MG}{c^2 d} = \frac{3MG}{c^2 a(1 - \epsilon^2)}} \quad - (35)$$

which is the planetary precession of the perihelion,

A.E.D.

6) The planetary precession is before the Thomas precession. Its effect is:

$$\theta \rightarrow \theta(1 + \alpha) \quad - (36)$$

where

$$\alpha = \gamma_T = \frac{dt}{d\tau} = \frac{3MG}{c^2 a} \quad - (37)$$

Therefore:

$$\boxed{dt \rightarrow \gamma_T d\tau} \quad - (38)$$

The precession changes to spin convention or angular velocity to:

$$\Omega = \frac{d\theta}{d\tau} = \gamma_T \frac{d\theta}{dt} \quad - (39)$$

This can be used as a definition of a relativistic angular velocity:

$$\Omega = \gamma_T \omega \quad - (40)$$

Conclusion The rotation of the Minkowski metric produces the universal planetary precession. It is not due to the incorrect Einstein equation.
